

# A GEOMETRIC INEQUALITY ON HYPERSURFACE IN HYPERBOLIC SPACE

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**ABSTRACT.** In this paper, we use the inverse curvature flow to prove a sharp geometric inequality on star-shaped and two-convex hypersurface in hyperbolic space.

## 1. INTRODUCTION

The classical Alexandrov-Fenchel inequalities for closed convex hypersurface  $\Sigma \subset \mathbb{R}^n$  state that

$$\int_{\Sigma} \sigma_m(\kappa) d\mu \geq C_{n,m} \left( \int_{\Sigma} \sigma_{m-1}(\kappa) d\mu \right)^{\frac{n-m-1}{n-m}}, \quad 1 \leq m \leq n-1 \quad (1)$$

where  $\sigma_m(\kappa)$  is the  $m$ -th elementary symmetric polynomial of the principal curvatures  $\kappa = (\kappa_1, \dots, \kappa_{n-1})$  of  $\Sigma$  and  $C_{n,m} = \frac{\sigma_m(1, \dots, 1)}{\sigma_{m-1}(1, \dots, 1)}$  is a constant. When  $m = 0$ , (1) is interpreted as the classical isoperimetric inequality

$$|\Sigma|^{\frac{1}{n-1}} \geq \bar{C}_n Vol(\Omega)^{\frac{1}{n}}, \quad (2)$$

which holds on all bounded domain  $\Omega \subset \mathbb{R}^n$  with boundary  $\Sigma = \partial\Omega$ . Here  $|\Sigma|$  is the area  $\Sigma$  and  $\bar{C}_n$  is a constant depending only on dimension  $n$ . Inequality (1) was generalized to star-shaped and  $m$ -convex hypersurface  $\Sigma \subset \mathbb{R}^n$  by Guan and Li [8] using the inverse curvature flow recently, where  $m$ -convex means that the principal curvature of  $\Sigma$  lies in the Garding's cone

$$\Gamma_m = \{\kappa \in \mathbb{R}^{n-1} | \sigma_i(\kappa) > 0, i = 1, \dots, m\}.$$

Recently, Huisken [11] showed that in the case  $m = 1$ , the assumption *star-shaped* can be replaced by *outward-minimizing*.

In this paper, we consider the hyperbolic space  $\mathbb{H}^n = \mathbb{R}^+ \times \mathbb{S}^{n-1}$  endowed with the metric

$$\bar{g} = dr^2 + \sinh^2 r g_{\mathbb{S}^{n-1}},$$

where  $g_{\mathbb{S}^{n-1}}$  is the standard round metric on the unit sphere  $\mathbb{S}^{n-1}$ . It's a natural question to establish some analogue inequalities of (1) for closed hypersurface in  $\mathbb{H}^n$ . In the case of  $m = 1$ ,  $\sigma_1 = \sigma_1(\kappa)$  is just the mean curvature  $H$  of  $\Sigma$ . Gallego and Solanes [6] have obtained a generalization of (1) to convex hypersurface in hyperbolic space using integral geometric methods,

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however, their result does not seem to be sharp. Denoting  $\lambda(r) = \sinh r$ , then  $\lambda'(r) = \cosh r$ . Recently, Brendle, Hung and Wang [3] proved the following inequality for star-shaped and mean convex (i.e.,  $H > 0$ ) hypersurface  $\Sigma \subset \mathbb{H}^n$ :

$$\int_{\Sigma} (\lambda' H - (n-1)\langle \bar{\nabla} \lambda', \nu \rangle) d\mu \geq (n-1)\omega_{n-1}^{\frac{1}{n-1}} |\Sigma|^{\frac{n-2}{n-1}} \quad (3)$$

where  $|\Sigma|$  is the area of  $\Sigma$  and  $\omega_{n-1}$  is the area of the unit sphere  $\mathbb{S}^{n-1} \subset \mathbb{R}^n$ . de Lima and Girao [4] also proved the following related inequality independently.

$$\int_{\Sigma} \lambda' H d\mu \geq (n-1)\omega_{n-1} \left( \left( \frac{|\Sigma|}{\omega_{n-1}} \right)^{\frac{n-2}{n-1}} + \left( \frac{|\Sigma|}{\omega_{n-1}} \right)^{\frac{n}{n-1}} \right), \quad (4)$$

Both inequalities (3) and (4) are sharp in the sense of that equality holds if and only if  $\Sigma$  is a geodesic sphere centered at the origin. Here, we say a closed hypersurface  $\Sigma \subset \mathbb{H}^n$  is star-shaped if the unit outward normal  $\nu$  satisfies  $\langle \nu, \partial_r \rangle > 0$  everywhere on  $\Sigma$ , which is also equivalent to that  $\Sigma$  can be parametrized by a graph

$$\Sigma = \{(r(\theta), \theta) | \theta \in \mathbb{S}^{n-1}\}$$

for some smooth function  $r$  on  $\mathbb{S}^{n-1}$ . We note that inequalities (3) and (4) have some applications in general relativity, see [3, 4, 14].

In this paper, we consider the case  $m = 2$ . We prove the following sharp inequality for star-shaped and two-convex hypersurface  $\Sigma \subset \mathbb{H}^n$ , where *two-convex* means that the principal curvature lies in the Garding's cone  $\Gamma_2$  everywhere on  $\Sigma$ .

**Theorem 1.** *If  $\Sigma \subset \mathbb{H}^n$  is a star-shaped and two-convex hypersurface, then*

$$\int_{\Sigma} \sigma_2 d\mu \geq \frac{(n-1)(n-2)}{2} \left( \omega_{n-1}^{\frac{2}{n-1}} |\Sigma|^{\frac{n-3}{n-1}} + |\Sigma| \right), \quad (5)$$

where  $\omega_{n-1}$  is the area of the unit sphere  $\mathbb{S}^{n-1} \subset \mathbb{R}^n$  and  $|\Sigma|$  is the area of  $\Sigma$ . The equality holds if and only if  $\Sigma$  is a geodesic sphere.

Note that there exists at least one elliptic point on a closed, connected hypersurface  $\Sigma$  in hyperbolic space  $\mathbb{H}^n$ . Proposition 3.2 in [1] shows that if  $\sigma_2$  is positive, then  $\sigma_1$  is automatically positive. So our assumption *two-convex* can also be replaced by  $\sigma_2 > 0$  on  $\Sigma$ .

The proof of Theorem 1 follows a similar argument as in [3, 4, 8]. We evolve  $\Sigma$  by a special case of the inverse curvature flow in [7], and consider the following quantity defined by

$$Q(t) = |\Sigma|^{-\frac{n-3}{n-1}} \left( \int_{\Sigma} \sigma_2 d\mu - \frac{(n-1)(n-2)}{2} |\Sigma| \right).$$

We show that  $Q(t)$  is monotone decreasing under the flow. Then we use the convergence result of the flow proved by Gerhardt to estimate a lower

bound of the limit of  $Q(t)$ :

$$\liminf_{t \rightarrow \infty} Q(t) \geq \frac{(n-1)(n-2)}{2} \omega_{n-1}^{\frac{2}{n-1}}.$$

In order to estimate this  $\liminf$ , we also use a sharp version Sobolev inequality on  $\mathbb{S}^{n-1}$  due to Beckner [2] as in [3]. Finally Theorem 1 follows easily from the monotonicity and the lower bound of  $\liminf_{t \rightarrow \infty} Q(t)$ .

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## 2. PRELIMINARIES

Let  $\Sigma \subset \mathbb{H}^n$  be a closed hypersurface with unit outward normal  $\nu$ . The second fundamental form  $h$  of  $\Sigma$  is defined by

$$h(X, Y) = \langle \bar{\nabla}_X \nu, Y \rangle$$

for any two tangent fields  $X, Y$ . The principal curvature  $\kappa = (\kappa_1, \dots, \kappa_{n-1})$  are the eigenvalues of  $h$  with respect to the induced metric  $g$  on  $\Sigma$ . For  $1 \leq m \leq n-1$ , the  $m$ -th elementary symmetric polynomial of  $\kappa$  is defined as

$$\sigma_m(\kappa) = \sum_{i_1 < i_2 < \dots < i_m} \kappa_{i_1} \cdots \kappa_{i_m},$$

which can also be viewed as function of the second fundamental form  $h_i^j = g^{jk} h_{ki}$ . In the sequel, we will simply write  $\sigma_m$  for  $\sigma_m(\kappa)$ . We first collect the following basic facts on  $\sigma_m$  (see, e.g, [9, 12, 13]):

**Lemma 2.** Denote  $(T_{m-1})_j^i = \frac{\partial \sigma_m}{\partial h_i^j}$  and  $(h^2)_i^j = g^{jl} g^{pk} h_{kl} h_{ip}$ . We have

$$\sum_{i,j} (T_{m-1})_j^i h_i^j = m \sigma_m, \tag{6}$$

$$\sum_{i,j} (T_{m-1})_j^i \delta_i^j = (n-m) \sigma_{m-1} \tag{7}$$

$$\sum_{i,j} (T_{m-1})_j^i (h^2)_i^j = \sigma_1 \sigma_m - (m+1) \sigma_{m+1} \tag{8}$$

Moreover, if  $\kappa \in \Gamma_m^+$ , we have the following Newton-MacLaurin inequalities

$$\frac{\sigma_{m-1} \sigma_{m+1}}{\sigma_m^2} \leq \frac{m(n-m-1)}{(m+1)(n-m)} \tag{9}$$

$$\frac{\sigma_1 \sigma_{m-1}}{\sigma_m} \geq \frac{m(n-1)}{n-m}, \tag{10}$$

and the equalities hold in (9),(10) at a given point if and only if  $\Sigma$  is umbilical there.

We now evolve  $\Sigma \subset \mathbb{H}^n$  by the following evolution equation

$$\partial_t X = F\nu, \quad (11)$$

where  $\nu$  is the unit outward normal to  $\Sigma_t = X(t, \cdot)$  and  $F$  is the speed function which may depend on the position vector, principal curvatures and time. Let  $g_{ij}$  be the induced metric and  $d\mu_t$  be its area element on  $\Sigma_t$ . We have the following evolution equations.

**Proposition 3.** *Under the flow (11), we have:*

$$\begin{aligned} \partial_t g_{ij} &= 2Fh_{ij} \\ \partial_t \nu &= -\nabla F, \\ \partial_t h_i^j &= -\nabla^j \nabla_i F - F(h^2)_i^j + F\delta_i^j, \\ \partial_t d\mu &= F\sigma_1 d\mu, \end{aligned} \quad (12)$$

$$\begin{aligned} \partial_t \sigma_m &= -\nabla^i ((T_{m-1})_i^j \nabla_j F) - F(\sigma_1 \sigma_m - (m+1)\sigma_{m+1}) \\ &\quad + (n-m)F\sigma_{m-1}, \end{aligned} \quad (13)$$

*Proof.* The first four equations follow from direct computations like in [10]. Now we calculate the evolution of  $\sigma_m$  (cf. [8])

$$\begin{aligned} \partial_t \sigma_m &= \frac{\partial \sigma_m}{\partial h_i^j} \partial_t h_i^j \\ &= -(T_{m-1})_j^i \nabla^j \nabla_i F - F(T_{m-1})_j^i (h^2)_i^j + F(T_{m-1})_j^i \delta_i^j \\ &= -\nabla^j ((T_{m-1})_j^i \nabla_i F) - F(\sigma_1 \sigma_m - (m+1)\sigma_{m+1}) \\ &\quad + (n-m)F\sigma_{m-1}, \end{aligned}$$

where in the last equality we used (7),(8) and the divergence free property of  $(T_{m-1})_j^i$  (see [13]).  $\square$

**Proposition 4.** *Under the flow (11), we have*

$$\frac{d}{dt} \int_{\Sigma} \sigma_m d\mu = (m+1) \int_{\Sigma} F\sigma_{m+1} d\mu + (n-m) \int_{\Sigma} F\sigma_{m-1} d\mu.$$

*Proof.* This proposition follows directly from (12), (13) and the divergence theorem.  $\square$

In [7] Gerhardt studied general inverse curvature flow of star-shaped hypersurface in hyperbolic space. For our purpose, we will use a special case of their result for the following flow

$$\partial_t X = \frac{n-2}{2(n-1)} \frac{\sigma_1}{\sigma_2} \nu. \quad (14)$$

**Theorem 5** (Gerhardt [7]). *If the initial hypersurface is star-shaped and strictly two-convex, then the solution for the flow (14) exists for all time  $t > 0$  and the flow hypersurfaces converge to infinity while maintaining star-shapedness and strictly two-convex. Moreover, the hypersurfaces become*

*strictly convex exponentially fast and more and more totally umbilical in the sense of*

$$|h_i^j - \delta_i^j| \leq Ce^{-\frac{t}{n-1}}, \quad t > 0,$$

*i.e., the principal curvatures are uniformly bounded and converge exponentially fast to one.*

### 3. MONOTONICITY

We define the quantity

$$Q(t) = |\Sigma_t|^{-\frac{n-3}{n-1}} \left( \int_{\Sigma_t} \sigma_2 d\mu - \frac{(n-1)(n-2)}{2} |\Sigma_t| \right),$$

where  $|\Sigma_t|$  is the area of  $\Sigma_t$ . In this section, we show that  $Q(t)$  is monotone decreasing along the flow (14).

**Proposition 6.** *Under the flow (14), the quantity  $Q(t)$  is monotone decreasing. Moreover,  $\frac{d}{dt}Q(t) = 0$  at some time  $t$  if and only if the surface  $\Sigma_t$  is totally umbilical.*

*Proof.* Under the flow (14), Proposition 4 and (12) imply that

$$\frac{d}{dt} \int_{\Sigma} \sigma_2 d\mu = \frac{3(n-2)}{2(n-1)} \int_{\Sigma} \frac{\sigma_1 \sigma_3}{\sigma_2} d\mu + \frac{(n-2)^2}{2(n-1)} \int_{\Sigma} \frac{\sigma_1^2}{\sigma_2} d\mu \quad (15)$$

$$\frac{d}{dt} |\Sigma_t| = \frac{(n-2)}{2(n-1)} \int_{\Sigma} \frac{\sigma_1^2}{\sigma_2} d\mu. \quad (16)$$

Combining (15), (16) and (9), we have

$$\frac{d}{dt} \left( \int_{\Sigma} \sigma_2 d\mu - (n-2) |\Sigma_t| \right) \leq \frac{n-3}{n-1} \int_{\Sigma} \sigma_2 d\mu. \quad (17)$$

By applying the Newton-MacLaurin inequality (10) in (16), we also have

$$\frac{d}{dt} |\Sigma_t| \geq |\Sigma_t|. \quad (18)$$

Then combining (17) and (18) gives that

$$\frac{d}{dt} \left( \int_{\Sigma} \sigma_2 d\mu - \frac{(n-1)(n-2)}{2} |\Sigma_t| \right) \leq \frac{n-3}{n-1} \left( \int_{\Sigma} \sigma_2 d\mu - \frac{(n-1)(n-2)}{2} |\Sigma_t| \right). \quad (19)$$

From Proposition 8 in the next section and (19), we know that the quantity

$$\int_{\Sigma} \sigma_2 d\mu - \frac{(n-1)(n-2)}{2} |\Sigma_t|$$

is nonnegative along the flow (14). Then inequalities (18) and (19) implies that

$$\frac{d}{dt} Q(t) \leq 0.$$

If the equality holds, the inequalities (9) and (10) assume equalities everywhere on  $\Sigma_t$ . Then  $\Sigma_t$  is totally umbilical.  $\square$

#### 4. THE ASYMPTOTIC BEHAVIOR OF MONOTONE QUANTITY

In this section, we use the convergence result of the flow (14) proved in [7] to estimate the lower bound of the limit of  $Q(t)$ . First we need the following sharp Sobolev inequality on  $\mathbb{S}^{n-1}$  ([2]).

**Lemma 7.** *For every positive function  $f$  on  $\mathbb{S}^{n-1}$ , we have*

$$\begin{aligned} & \int_{\mathbb{S}^{n-1}} f^{n-3} dvol_{\mathbb{S}^{n-1}} + \frac{n-3}{n-1} \int_{\mathbb{S}^{n-1}} f^{n-5} |\nabla f|^2 dvol_{\mathbb{S}^{n-1}} \\ & \geq \omega_{n-1}^{\frac{2}{n-1}} \left( \int_{\mathbb{S}^{n-1}} f^{n-1} dvol_{\mathbb{S}^{n-1}} \right)^{\frac{n-3}{n-1}}. \end{aligned}$$

Moreover, equality holds if and only if  $f$  is a constant.

*Proof.* From Theorem 4 in [2], for any positive smooth function  $w$  on  $\mathbb{S}^{n-1}$ , we have the following inequality

$$\begin{aligned} & \frac{4}{(n-1)(n-3)} \int_{\mathbb{S}^{n-1}} |\nabla w|^2 dvol_{\mathbb{S}^{n-1}} + \int_{\mathbb{S}^{n-1}} w^2 dvol_{\mathbb{S}^{n-1}} \\ & \geq \omega_{n-1}^{\frac{2}{n-1}} \left( \int_{\mathbb{S}^{n-1}} w^{\frac{2(n-1)}{n-3}} dvol_{\mathbb{S}^{n-1}} \right)^{\frac{n-3}{n-1}}. \end{aligned}$$

Moreover equality holds if and only if  $w$  is constant. For any positive function  $f$  on  $\mathbb{S}^{n-1}$ , by letting  $w = f^{\frac{n-3}{2}}$ , we have

$$\begin{aligned} & \int_{\mathbb{S}^{n-1}} f^{n-3} dvol_{\mathbb{S}^{n-1}} + \frac{n-3}{n-1} \int_{\mathbb{S}^{n-1}} f^{n-5} |\nabla f|^2 dvol_{\mathbb{S}^{n-1}} \\ & \geq \omega_{n-1}^{\frac{2}{n-1}} \left( \int_{\mathbb{S}^{n-1}} f^{n-1} dvol_{\mathbb{S}^{n-1}} \right)^{\frac{n-3}{n-1}} \end{aligned}$$

and equality holds if and only if  $f$  is a constant.  $\square$

**Proposition 8.** *Under the flow (14), we have*

$$\liminf_{t \rightarrow \infty} Q(t) \geq \frac{(n-1)(n-2)}{2} \omega_{n-1}^{\frac{2}{n-1}}. \quad (20)$$

*Proof.* Recall that star-shaped hypersurfaces can be written as graphs of function  $r = r(t, \theta), \theta \in \mathbb{S}^{n-1}$ . Denote  $\lambda(r) = \sinh(r)$ , then  $\lambda'(r) = \cosh(r)$ . We next define a function  $\varphi$  on  $\mathbb{S}^{n-1}$  by  $\varphi(\theta) = \Phi(r(\theta))$ , where  $\Phi(r)$  is a positive function satisfying  $\Phi' = 1/\lambda$ . Let  $\theta = \{\theta^j\}, j = 1, \dots, n-1$  be a coordinate system on  $\mathbb{S}^{n-1}$  and  $\varphi_i, \varphi_{ij}$  be the covariant derivatives of  $\varphi$  with respect to the metric  $g_{\mathbb{S}^{n-1}}$ . Define

$$v = \sqrt{1 + |\nabla \varphi|_{\mathbb{S}^{n-1}}^2}.$$

From [7], we know that

$$\lambda = O(e^{\frac{t}{n-1}}), \quad |\nabla \varphi|_{\mathbb{S}^{n-1}} + |\nabla^2 \varphi|_{\mathbb{S}^{n-1}} = O(e^{-\frac{t}{n-1}}) \quad (21)$$

Since  $\lambda' = \sqrt{1 + \lambda^2}$ , we have

$$\lambda' = \lambda(1 + \frac{1}{2}\lambda^{-2} + O(e^{-\frac{4t}{n-1}})) \quad (22)$$

From (21) we also have

$$\frac{1}{v} = 1 - \frac{1}{2}|\nabla\varphi|_{\mathbb{S}^{n-1}}^2 + O(e^{-\frac{4t}{n-1}}) \quad (23)$$

In terms of  $\varphi$ , we can express the metric and second fundamental form of  $\Sigma$  as following (see, e.g, [3, 5])

$$\begin{aligned} g_{ij} &= \lambda^2(\sigma_{ij} + \varphi_i \varphi_j) \\ h_{ij} &= \frac{\lambda'}{v\lambda}g_{ij} - \frac{\lambda}{v}\varphi_{ij}, \end{aligned}$$

where  $\sigma_{ij} = g_{\mathbb{S}^{n-1}}(\partial_{\theta^i}, \partial_{\theta^j})$ . Denote  $a_i = \sum_k \sigma^{ik} \varphi_{ki}$  and note that  $\sum_i a_i = \Delta_{\mathbb{S}^{n-1}}\varphi$ . By (21), the principal curvatures of  $\Sigma_t$  has the following form

$$\kappa_i = \frac{\lambda'}{v\lambda} - \frac{a_i}{v\lambda} + O(e^{-\frac{4t}{n-1}}), \quad i = 1, \dots, n-1.$$

Then we have

$$\begin{aligned} \sigma_2 &= \sum_{i < j} \kappa_i \kappa_j \\ &= \frac{(n-1)(n-2)}{2} \left( \frac{\lambda'}{v\lambda} \right)^2 - (n-2) \frac{\lambda' \Delta_{\mathbb{S}^{n-1}}\varphi}{v^2 \lambda^2} + O(e^{-\frac{4t}{n-1}}). \end{aligned}$$

By using (22) and (23),

$$\begin{aligned} \sigma_2 &= \frac{(n-1)(n-2)}{2} (1 + \lambda^{-2} - |\nabla\varphi|_{\mathbb{S}^{n-1}}^2) \\ &\quad - (n-2)\lambda^{-1} \Delta_{\mathbb{S}^{n-1}}\varphi + O(e^{-\frac{4t}{n-1}}). \end{aligned}$$

On the other hand,

$$\sqrt{\det g} = (\lambda^{n-3} + O(e^{\frac{(n-3)t}{n-1}})) \sqrt{\det g_{\mathbb{S}^{n-1}}}.$$

So we have

$$\begin{aligned}
& \int_{\Sigma_t} \sigma_2 d\mu - \frac{(n-1)(n-2)}{2} |\Sigma_t| \\
&= \int_{\mathbb{S}^{n-1}} \lambda^{n-1} (\sigma_2 - \frac{(n-1)(n-2)}{2}) dvol_{\mathbb{S}^{n-1}} + O(e^{\frac{(n-5)t}{n-1}}) \\
&= \frac{(n-1)(n-2)}{2} \int_{\mathbb{S}^{n-1}} (\lambda^{n-3} - \lambda^{n-1} |\nabla \varphi|_{\mathbb{S}^{n-1}}^2) dvol_{\mathbb{S}^{n-1}} \\
&\quad - (n-2) \int_{\mathbb{S}^{n-1}} \lambda^{n-2} \Delta_{\mathbb{S}^{n-1}} \varphi dvol_{\mathbb{S}^{n-1}} + O(e^{\frac{(n-5)t}{n-1}}) \\
&= \frac{(n-1)(n-2)}{2} \int_{\mathbb{S}^{n-1}} (\lambda^{n-3} - \lambda^{n-1} |\nabla \varphi|_{\mathbb{S}^{n-1}}^2) dvol_{\mathbb{S}^{n-1}} \\
&\quad + (n-2)^2 \int_{\mathbb{S}^{n-1}} \lambda^{n-3} \nabla \lambda \nabla \varphi dvol_{\mathbb{S}^{n-1}} + O(e^{\frac{(n-5)t}{n-1}}).
\end{aligned}$$

Since  $\nabla \lambda = \lambda \lambda' \nabla \varphi$ , by using (22), we deduce that

$$\begin{aligned}
& \int_{\Sigma_t} \sigma_2 d\mu - \frac{(n-1)(n-2)}{2} |\Sigma_t| \\
&= \frac{(n-1)(n-2)}{2} \int_{\mathbb{S}^{n-1}} (\lambda^{n-3} + \frac{n-3}{n-1} \lambda^{n-5} |\nabla \lambda|^2) dvol_{\mathbb{S}^{n-1}} + O(e^{\frac{(n-5)t}{n-1}}).
\end{aligned} \tag{24}$$

Moreover,

$$|\Sigma_t|^{\frac{n-3}{n-1}} = \left( \int_{\mathbb{S}^{n-1}} \lambda^{n-1} dvol_{\mathbb{S}^{n-1}} \right)^{\frac{n-3}{n-1}} + O(e^{\frac{(n-5)t}{n-1}}). \tag{25}$$

Using Lemma 7, we can complete the proof of Proposition 8 by combining (24) and (25).  $\square$

We now complete the proof of Theorem 1

*Proof of Theorem 1.* Since  $Q(t)$  is monotone decreasing, we have

$$Q(0) \geq \liminf_{t \rightarrow \infty} Q(t) \geq \frac{(n-1)(n-2)}{2} \omega_{n-1}^{\frac{2}{n-1}}.$$

This gives that the initial hypersurface  $\Sigma$  satisfies

$$\left( \int_{\Sigma} \sigma_2 d\mu - \frac{(n-1)(n-2)}{2} |\Sigma| \right) \geq \frac{(n-1)(n-2)}{2} \omega_{n-1}^{\frac{2}{n-1}} |\Sigma|^{\frac{n-3}{n-1}},$$

which is equivalent to the inequality (5) in Theorem 1. Now we assume that equality holds in (5), which implies that  $Q(t)$  is a constant. Then Proposition 6 implies  $\Sigma_t$  is umbilical and therefore a geodesic sphere. It is also easy to see that if  $\Sigma$  is a geodesic sphere of radius  $r$ , then the area of  $\Sigma$

is  $|\Sigma| = \omega_{n-1} \sinh^{n-1} r$  and the integral of  $\sigma_2$  is

$$\begin{aligned}\int_{\Sigma} \sigma_2 d\mu &= \frac{(n-1)(n-2)}{2} \omega_{n-1} \coth^2 r \sinh^{n-1} r \\ &= \frac{(n-1)(n-2)}{2} \omega_{n-1} (\sinh^{n-1} r + \sinh^{n-3} r) \\ &= \frac{(n-1)(n-2)}{2} \left( |\Sigma| + \omega_{n-1}^{\frac{2}{n-1}} |\Sigma|^{\frac{n-3}{n-1}} \right).\end{aligned}$$

Hence the equality holds in (5) on a geodesic sphere. This completes the proof of Theorem 1.  $\square$

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